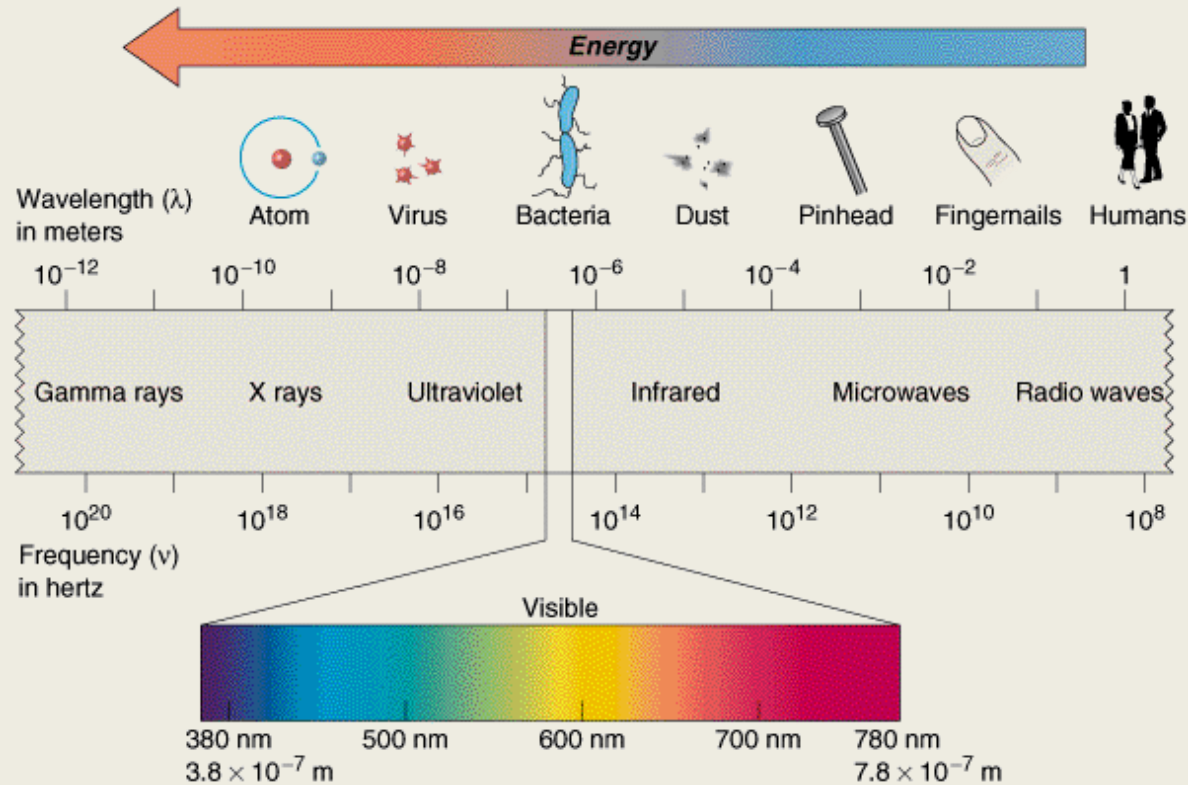


AH Chemistry – Unit 1

Electromagnetic Radiation and
Atomic Spectra

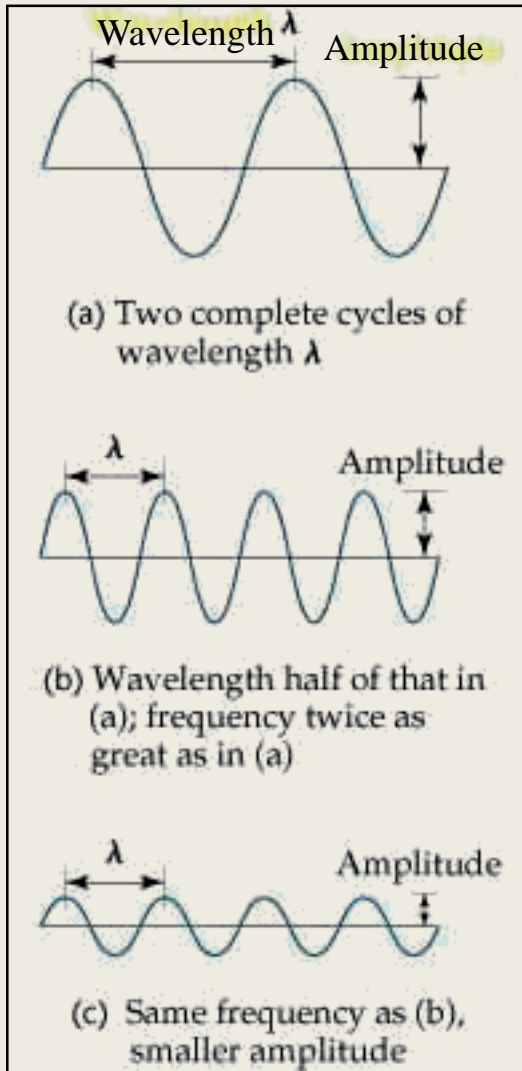
Introduction



This topic describes the ***Electromagnetic Spectrum*** and how it can interact with atoms, ***Spectroscopy***.

Much information about ***Electronic Structure*** comes from spectroscopic evidence.

The Wave Nature of Light

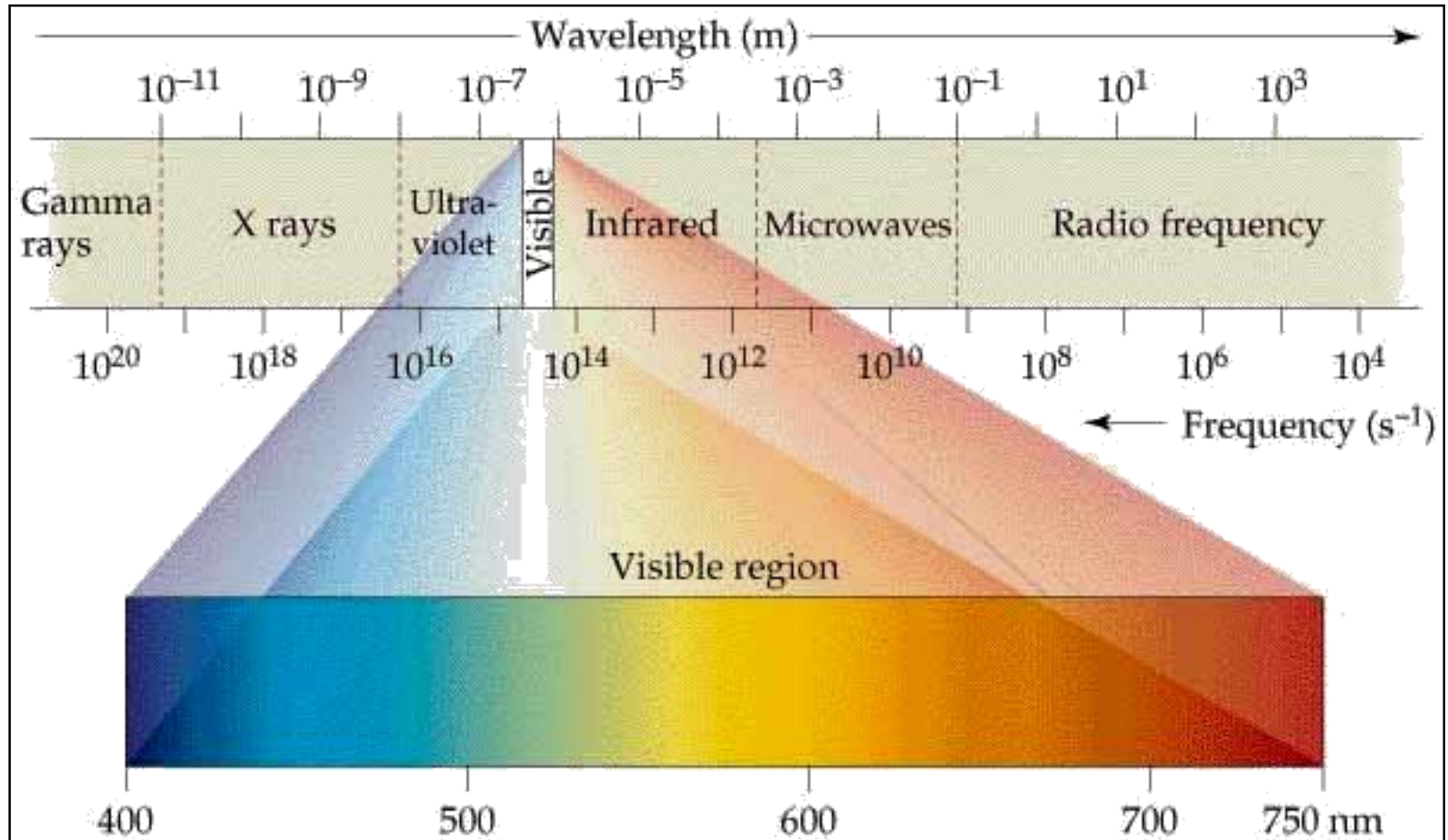


- All waves have a characteristic **wavelength**, λ , measured in metres (m) to nanometres (nm)
- The **frequency**, ν , of a wave is the number of waves which pass a point in one second measured in Hertz (Hz) or per seconds (s^{-1})
- The **speed** of a wave, c , is given by its frequency multiplied by its wavelength:

$$c = \nu\lambda$$

- For light, $c = 3.00 \times 10^8 \text{ ms}^{-1}$
- Another unit of ‘frequency’ used in spectroscopy is the **wavenumber** ($1/\lambda$), **ν** , measured in m^{-1} or cm^{-1}

Electromagnetic Radiation



The Particle Nature of Light 1

Planck: energy can only be absorbed or released from atoms in certain amounts called **quanta**

To understand **quantization**, consider the notes produced by a violin (**continuous**) and a piano (**quantized**):

*a violin can produce **any note** by placing the fingers at an appropriate spot on the fingerboard.*

*A piano can only produce **certain notes** corresponding to the keys on the keyboard.*

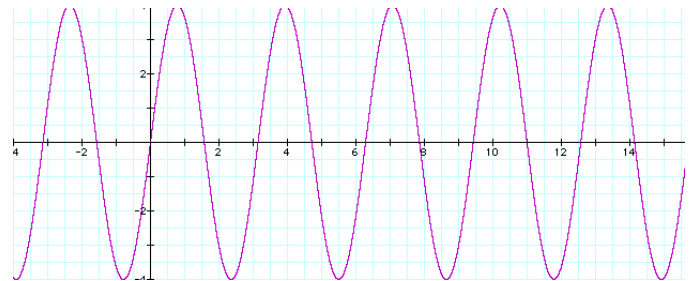
The Particle Nature of Light 2

Electromagnetic Radiation can also be thought of as a stream of very small particles known as **photons**

Electromagnetic Radiation exhibits wave-particle **dual properties**.

The **energy** (**E**) of a **photon** (particle) is related to the **frequency** (wave) of the radiation as follows:

$$E = h\nu$$



where h is **Planck's constant** (6.63×10^{-34} J s).

Energy Calculations 1

$$E = h\nu \quad \text{or} \quad E = hc / \lambda$$

The **energy** calculated would be in Joules (J) and would be a very small quantity.

Normally, we would calculate the energy transferred by the **emission** or **absorption** of **one mole of photons** as follows:

$$E = Lh\nu \quad \text{or} \quad E = Lhc / \lambda$$

Where L is the **Avogadro Constant**, 6.02×10^{23} and E would now be in J mol^{-1} or kJ mol^{-1} .

Energy Calculations 2

For example, a neon strip light emitted light with a wavelength of 640 nm.

$$640 \text{ nm} = 640 \times 10^{-9} \text{ m} = 6.40 \times 10^{-7} \text{ m}.$$

For each *photon*:

$$\begin{aligned} E &= h c / \lambda \\ &= 6.63 \times 10^{-34} \times 3.00 \times 10^8 / 6.40 \times 10^{-7} \\ &= 3.11 \times 10^{-19} \text{ J} \end{aligned}$$

For *1 mole of photons*:

$$\begin{aligned} E &= 3.11 \times 10^{-19} \times 6.02 \times 10^{23} \text{ J} \\ &= 1.87 \times 10^5 \text{ J mol}^{-1} \\ &= 187 \text{ kJ mol}^{-1} \end{aligned}$$

Atomic Spectra

Atomic emission spectra provided significant contributions to the modern picture of atomic structure

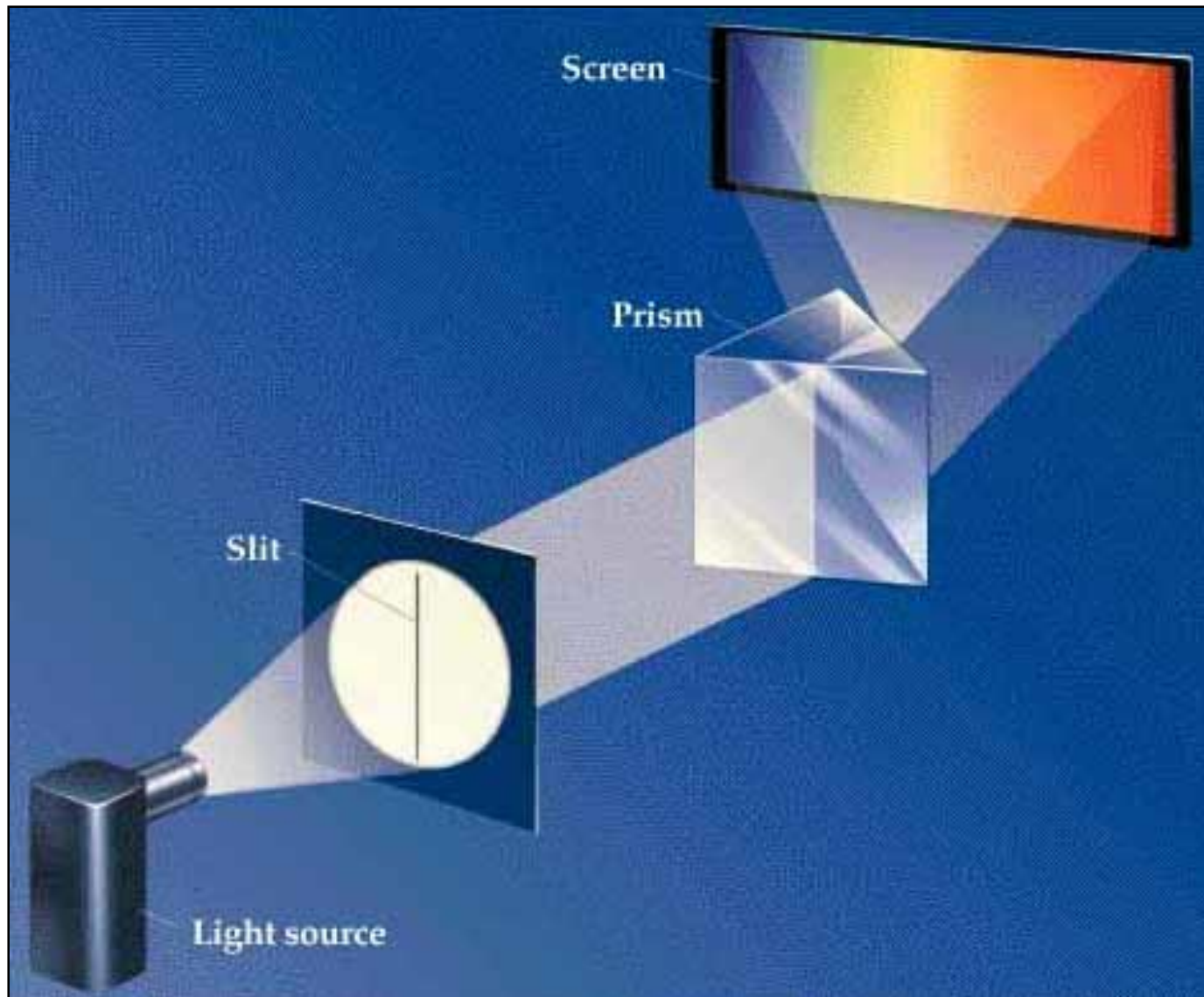


Radiation composed of only one wavelength is called ***monochromatic***

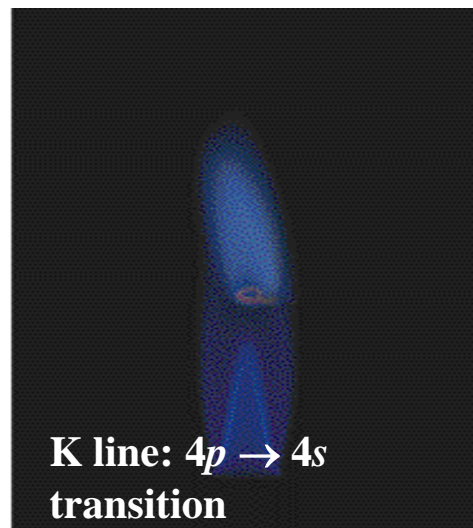
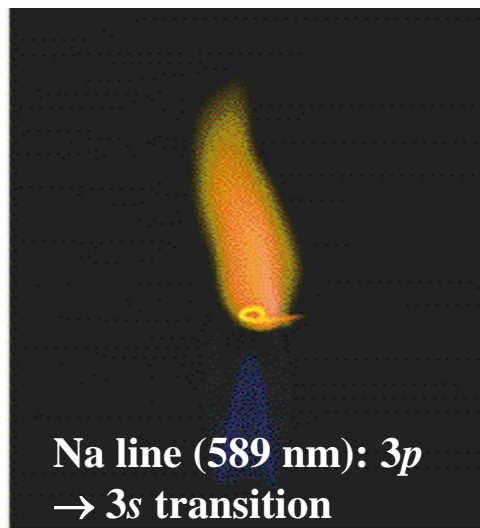
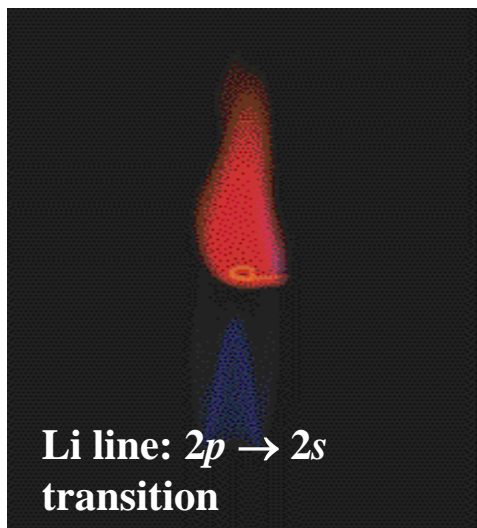
Radiation that spans a whole array of different wavelengths is called ***continuous***

White light can be separated into a continuous ***spectrum*** of colors.

Atomic Spectra 2



Emission Spectroscopy

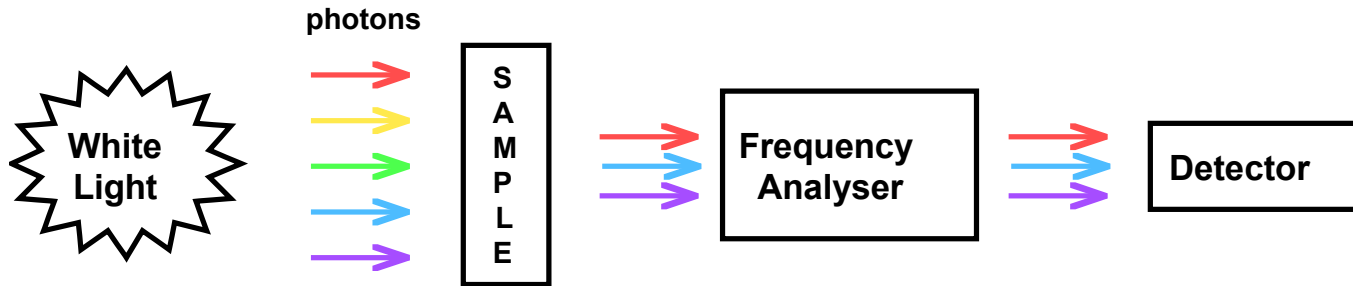


In ***emission spectroscopy***, light of certain wavelengths is emitted as ‘excited’ electrons drop down from higher energy levels.

The spectrum is examined to see the wavelengths emitted.

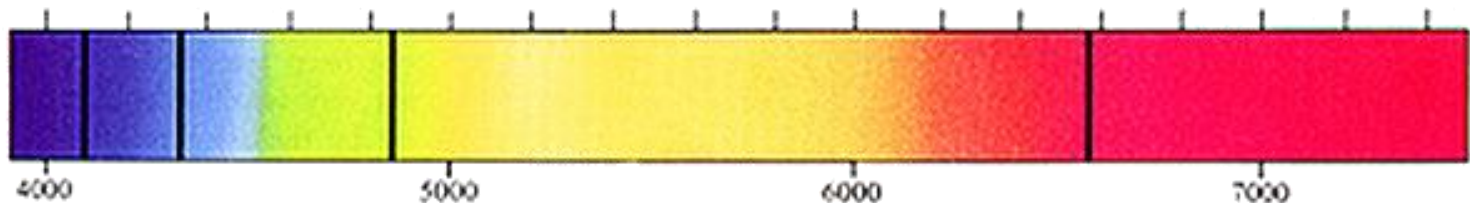


Absorption Spectroscopy



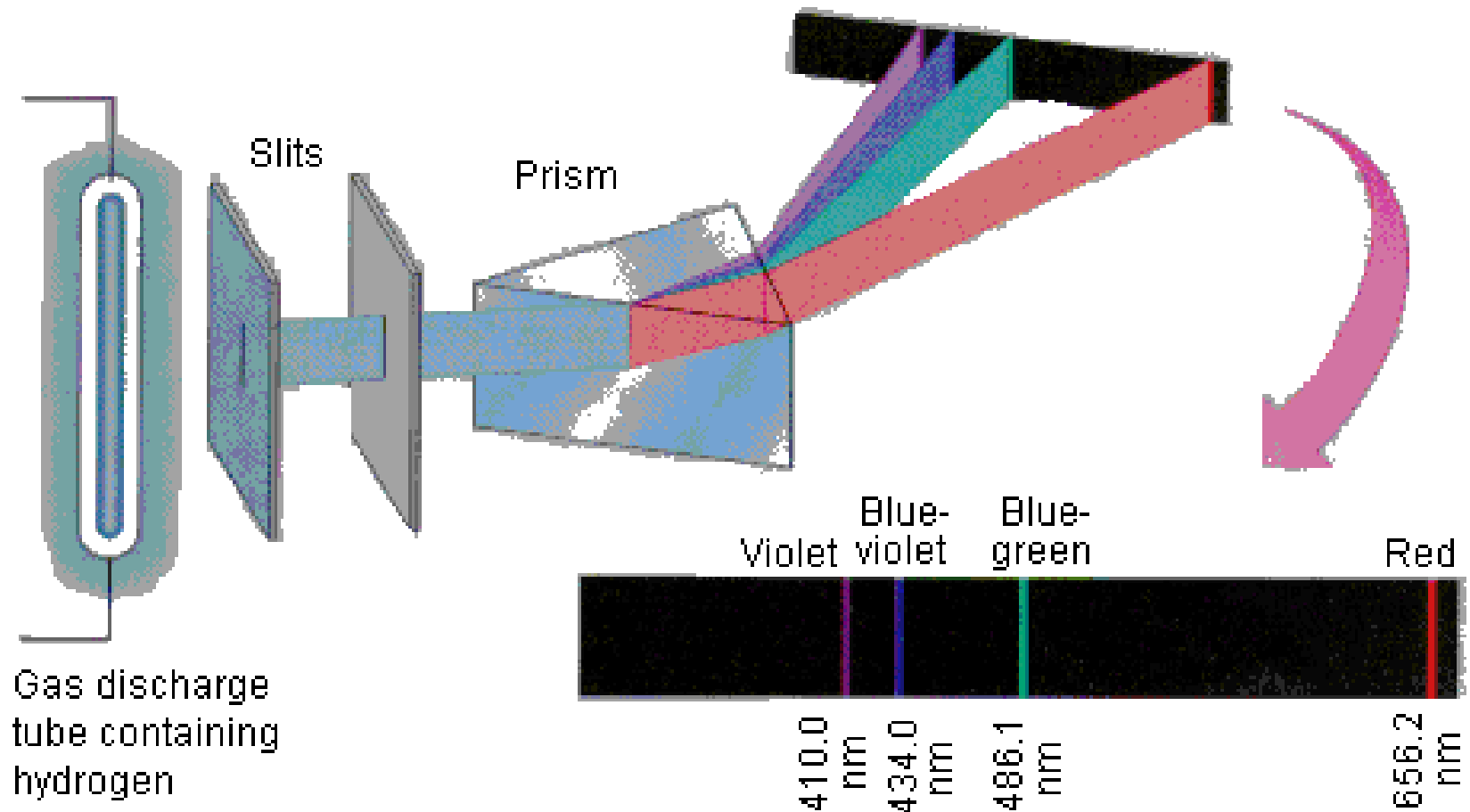
In **absorption spectroscopy**, light of certain wavelengths is absorbed and electrons are promoted to higher energy levels.

The spectrum is examined to see which wavelengths have been absorbed.



The **intensity** of the light absorbed is proportional to the **quantity** of the atoms/ions in the sample. Calibration samples can be prepared, intensities measured and unknown concentrations determined (PPA)

Hydrogen Emission Spectrum



Line Spectra

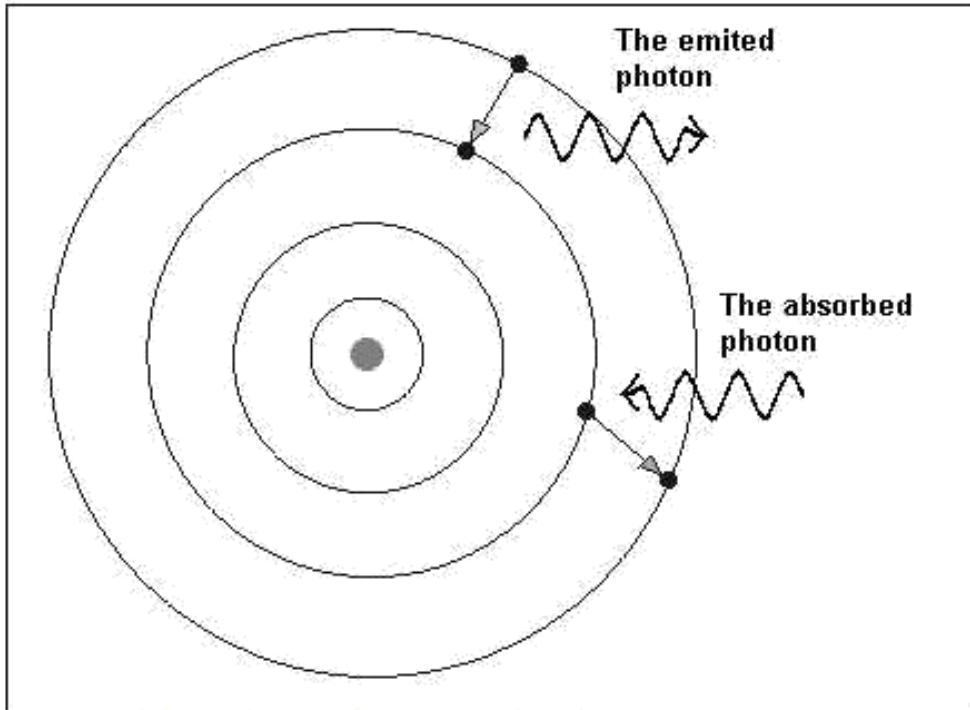
The spectrum obtained when hydrogen atoms are **excited** shows four lines: red, blue-green, blue and indigo



Bohr deduced that the colours were due to the movement of **electrons** from a higher energy level back to the ‘ground state’.

The significance of a Line Spectrum is that it suggests that electrons can only occupy certain **fixed energy levels**

Bohr's Model



The electron emits or absorbs the energy changing the orbits.

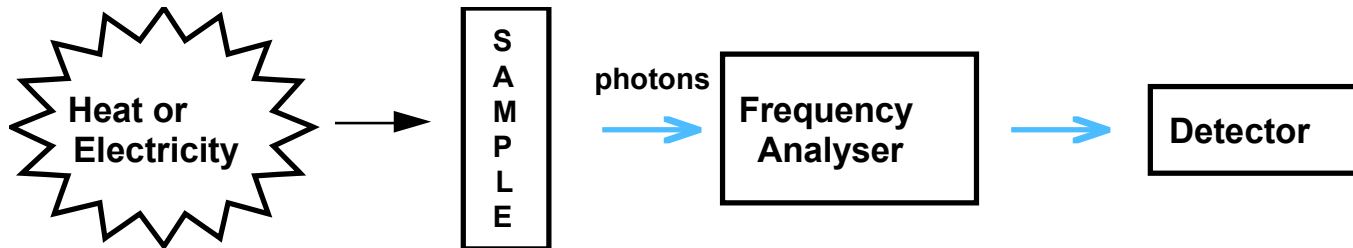
These fixed energy levels are what we have always called our ***electron shells***

A photon of light is emitted or absorbed when an electron changes from one energy level (shell) to another.

Emission Spectroscopy 2

Each element provides a ***characteristic spectrum*** which can be used to ***identify*** the element. Analysing light from stars etc, tell us a lot about the elements present.

The ***intensity*** of a particularly strong line in an element's spectrum can be measured. The ***intensity*** of the light emitted is proportional to the ***quantity*** of the atoms/ions in the sample. ***Calibration samples*** can be prepared, intensities measured and unknown concentrations determined. ***Analytical tool.***

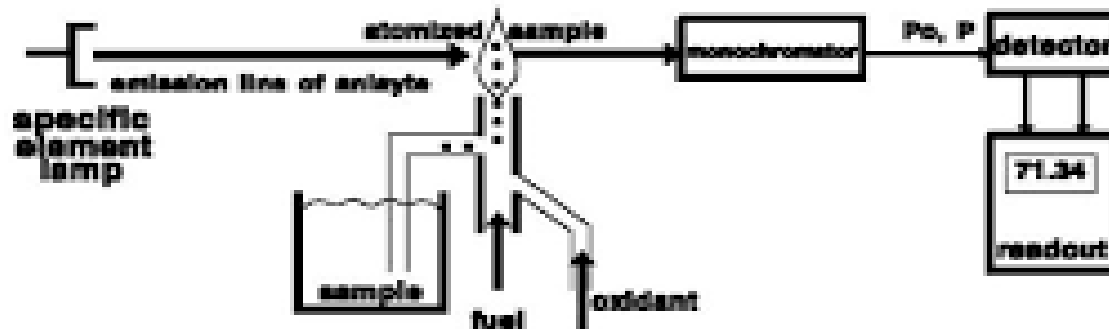


How much of an element?

If a blank or reference sample is introduced, it is possible to quantify how much of an element is present in a sample, as the intensity of light absorbed will be proportional to concentration.

Atomic Spectroscopy with Flames

Atomic Absorption Spectroscopy



$$A \propto C$$

P_0 = light intensity w/ blank
 P = light intensity w/ sample
 $A = \log[P_0/P] = kbC$
 b = flame path; C = sample concn
 k depends on absorptivity and flow